

Spectral Compression of Charged Black Holes: (Q) for Kerr-Newman

How Electric Charge Modifies the Analyticity Parameter

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Abstract

The quasinormal mode analyticity parameter $\nu = \ell / \ell_0$ is universal for Kerr (uncharged) black holes [Paper I]. We extend this analysis to Kerr-Newman spacetimes, computing $\nu(a, Q)$ across the full parameter space. The WKB approximation predicts $\nu \rightarrow 3$ whenever the standard light ring structure persists. We map the (a, Q) surface and identify [potential deviations / continued universality] near the extremal Kerr-Newman limit where $a^2 + Q^2 \rightarrow M^2$.

1. Introduction

1.1 Motivation

- Paper I: $\nu = 3$ for Kerr ($Q = 0$), universal across spins
- Natural question: what happens with charge $Q > 0$?
- Relevant for: primordial black holes, dark photon models, charged BH solutions in modified gravity

1.2 Kerr-Newman QNMs

- Teukolsky equation generalizes to KN with additional Q -dependent terms
- QNM frequencies $\omega(M, a, Q)$ computed by various methods
- Light ring at $r_{\text{ph}}(a, Q)$ — shifts with charge

2. Method

2.1 QNM Computation

- Use Leaver's continued fraction for KN (Berti & Stergioulas tables)
- Or: direct computation via spectral methods
- Compute $\ell=2, m=2$ gravitational QNMs for grid: $a/M \in [0, 0.99], Q/M \in [0, 0.99]$
- Subject to extremality bound $a^2 + Q^2 \leq M^2$

2.2 WKB Analysis

- Light ring radius $r_{\text{ph}}(a, Q)$
- Surface gravity $\kappa(a, Q)$
- WKB prediction: $\nu \rightarrow (n + 1/2)/(n + 1/2) = 3$

- Does this persist for all Q ?

2.3 Surface

- Compute (a, Q) on 50×50 grid
- Identify: contour at $= 3$, maximum/minimum, gradients

3. Results

3.1 The (a, Q) Surface

[Heatmap: a vs Q , color =] [Prediction: $= 3$ everywhere except possibly near extremal KN]

3.2 WKB Universality

- If $= 3$ for all $(a, Q) \rightarrow$ “super-universality”, $= 3$ is a topological invariant of ANY stationary BH light ring
- If (Q) deviates \rightarrow charge provides a new handle

3.3 Extremal Limit

- At $a^2 + Q^2 \rightarrow M^2$: what happens to $?$
- Kerr limit ($Q \rightarrow 0$): $\rightarrow 3.00$
- Reissner-Nordström limit ($a \rightarrow 0$): $(Q) = ?$

4. Implications

4.1 Charged BH Spectroscopy

- If $?$ depends on Q : measuring $?$ constrains Q
- Combined with mass/spin from fundamental mode \rightarrow three independent tests

4.2 No-Hair Theorem Test

- For Kerr ($Q = 0$): $= 3$ is a prediction
- Any deviation from $= 3$ in an observed ringdown \rightarrow evidence for charge or beyond-GR physics

4.3 Dark Sector

- Millicharged BHs: small $Q \rightarrow$ small $?$ deviation
- Observable threshold: how much Q is needed for $\Delta >$ measurement error?

5. Conclusion

[Map of (a, Q) provides theoretical predictions. If universality persists $\rightarrow = 3$ is a fundamental property of ALL BH light rings.]

References

[TBD — Berti et al. 2009, Kokkotas & Schutz 1999, Dias et al. for KN QNMs, Paper I]