

The Analyticity Parameter of Neutron Star Post-Merger Oscillations

Is $\gamma = 3$ Unique to Black Holes?

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Scaffold

Abstract

For black hole ringdown, the analyticity parameter $\gamma = \omega / \text{Im}(\omega)$ is universal — independent of mass, spin, and multipole [Paper I]. This universality originates from the topological structure of the photon sphere via the WKB approximation. Neutron star post-merger oscillations lack a photon sphere; their oscillation spectra depend on the nuclear equation of state (EOS). We compute γ_{NS} from numerical relativity simulations of binary neutron star mergers across multiple EOS models and show that γ_{NS} is EOS-dependent, ranging from [TBD] to [TBD]. This makes γ_{NS} a novel EOS discriminant, complementing the tidal deformability Λ and the dominant post-merger frequency f .

1. Introduction

[Paper I established $\gamma = 3$ for BH QNMs. Key question: does this extend to NS?]

1.1 Why γ_{NS} Should Differ

- No light ring \rightarrow WKB argument fails
- Oscillation modes (f, p, w, g) have EOS-dependent structure
- The damping rates depend on the compactness, the crust, viscosity
- γ_{NS} encodes structural information about the remnant

1.2 NS Post-Merger Spectrum

- Dominant f mode ($\sim 2-4$ kHz)
- Secondary peaks: f_p , f_g , spiral-arm frequencies
- w-modes (spacetime modes) at higher frequencies
- The ratio of damping rates between these $\rightarrow \gamma_{\text{NS}}$

2. Method

2.1 NR Waveforms

- Source: CoRe database (Dietrich et al.)
- Select 4-6 simulations spanning soft/stiff EOS: SLy, APR4, H4, MS1, DD2
- Extract post-merger GW signal (after merger, $t > t_{\text{merger}}$)
- Focus on the dominant (2,2) mode

2.2 Mode Extraction

- Prony/matrix pencil method for multi-frequency extraction
- Identify f (fundamental) and overtones/secondary frequencies
- Compute τ_n for each identified mode
- Define $\tau_{NS} = \tau_1 / \tau_2$ (ratio of first secondary to dominant damping rate)

2.3 Comparison Framework

- BH control: compute τ_{BH} from same-mass BH QNMs
- Plot $\tau_{NS}(\text{EOS})$ vs $\tau_{BH} = 3$

3. Results

3.1 τ_{NS} by EOS

[Table: EOS | f (kHz) | τ_n | τ_{NS}] [Prediction: τ_{NS} varies with EOS, potentially 1.5-5]

3.2 τ_{NS} as EOS Discriminant

- Correlation with compactness $C = M/R$
- Correlation with tidal deformability Λ
- Information content: does τ_{NS} add to Λ ?

3.3 Comparison with BH Universality

- $\tau_{BH} = 3$ (universal)
- $\tau_{NS} = ?$ (EOS-dependent)
- The difference fingerprints the object: BH vs NS from τ alone

4. Discussion

4.1 Observational Prospects

- Post-merger signal detectable with ET/CE at ~ 10 Mpc
- τ_{NS} measurable if $\text{SNR} > \sim 20$ in post-merger band
- Complementary to pre-merger tidal constraints

4.2 BH Mimickers

- If a compact object has $\tau = 3$ in ringdown \rightarrow strong evidence for BH
- If $\tau \neq 3 \rightarrow$ evidence for exotic compact object or NS remnant

5. Conclusion

[τ_{NS} is EOS-dependent, making it a new observable. Combined with f and Λ , it provides a three-pronged constraint on the nuclear EOS.]

References

[TBD — key refs: Dietrich et al. CoRe database, Bauswein et al. f universal relations, Paper I]